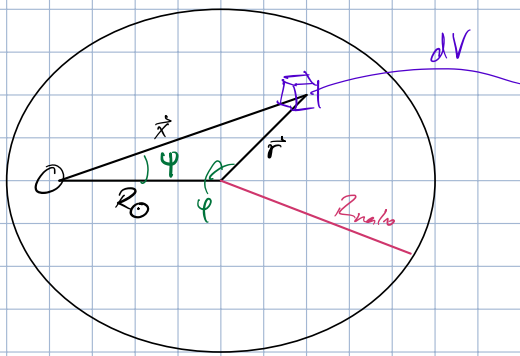


iAMUH!

Indirect Detection



DM may decay or annihilate

decay must be super rare:  $\tau_{\alpha}^{\text{decay}} \gg 10 \text{ Gyr}$

$$\tau_{\alpha} = (R_{\alpha})^{-1}$$

↳ decay rate

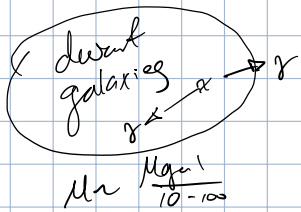
$$\chi \text{ --- } \left\langle \frac{A}{A} \right\rangle$$

time →

$$\chi \rightarrow e^+e^-$$

$$\sim \gamma\gamma \quad \left( E_{\gamma} = \frac{m_{\chi}}{2} \right)$$

isotropic!



Mass distribution:

$$\rho_{\chi}(r) = \frac{4 f_s}{\left(\frac{r}{r_s}\right) \left(1 + \frac{r}{r_s}\right)^2}$$

$r_s$  - scale radius  $\sim 10 \text{ kpc}$   
 $f_s$  - scale density

what is flux from  $\chi \rightarrow \gamma\gamma\gamma\gamma\dots$  decay

for each decay channel  $\chi \rightarrow FF \rightarrow \frac{dN_{\chi}}{dE_{\chi}}$ , model dependent  $\gamma$ -distribution

flux in detector from  $dV$ ?

$$\frac{dN_{\chi}^{\text{obs}}}{dA_{\text{det}}} = \frac{1}{4\pi r^2} \underbrace{dN_{\chi}}_{\substack{\# \chi \text{ particles in box} \\ dV}} \underbrace{P_{\text{decay}}}_{\substack{\text{probability of decay}}}$$

$$dN_{\chi} \equiv n_{\chi}(r(\vec{x})) dV = \frac{\rho_{\chi}(r(\vec{x}))}{m_{\chi}} r^2 \sin\theta dr d\Omega$$

$$P_{\text{decay}} = R_{\chi} \cdot dt = \frac{dR_{\chi}}{dE_{\chi}} \cdot dE_{\chi} \cdot dt$$

decay rate into photons  $\gamma$

$$\frac{dN_r^{\text{obs}}}{dA dt} = \frac{n[\vec{r}(\vec{x})]}{4\pi x^2} \cdot x^2 dx d\Omega_\psi \cdot \frac{dR_\alpha}{dE_\gamma} \cdot dE_\gamma \cdot dt$$

$$\frac{dN_r^{\text{obs}}}{dA dt \cdot dt \cdot dE_\gamma} = \frac{n[\vec{r}(\vec{x})]}{4\pi} \cdot dx d\Omega_\psi \left( \frac{dR_\alpha}{dE_\gamma} \right)$$

$$\Phi = \text{flux} \sim \frac{1}{\text{area} \cdot \text{time}}$$

integrate over all lines of sight

$$\frac{d\Phi_r}{dE_\gamma} = \frac{n[\vec{r}(\vec{x})]}{4\pi} \cdot dx d\Omega_\psi \left( \frac{dR_\alpha}{dE_\gamma} \right)$$

add all box contributions along line of sight  $\vec{x}$

$$\frac{d\Phi_r}{dE_\gamma} = \frac{1}{4\pi} \frac{dR_\alpha}{dE_\gamma} \cdot \int d\Omega_\psi \int d^3\vec{x} \cdot \frac{\rho_\alpha[\vec{r}(x)]}{m_\alpha}$$

whats  $\vec{r}(x)$ ?  $\vec{r}^2 = \vec{x}^2 + \vec{R}_0^2 - 2\vec{x} \cdot \vec{R}_0 \cos\psi$

maximum  $x$ :  $x_{\text{max}}^2 = R_0^2 + R_{\text{halo}}^2 - 2R_0 R_{\text{halo}} \cos\psi$

$$\frac{d\Phi_r}{dE_\gamma} = \frac{1}{4\pi} \frac{dR_\alpha}{dE_\gamma} \int d\Omega_\psi \cdot \int_0^{x_{\text{max}}(\psi)} \frac{\rho_\alpha[\vec{r}(x)]}{m_\alpha} dx$$

@  $x_{\text{max}}(\psi)$   $|\vec{r}| = R_{\text{halo}} \sim 200 \text{ kpc}$

$$d\Omega_\psi = d(\cos\psi)$$

$$x_{\text{max}}^2 = R_0^2 + R_H^2 - 2R_H \cdot R_0 \cos\psi$$

$$\int d\Omega_\psi = \sin\psi d\psi d\varphi$$

$$\frac{d\Phi_\gamma}{dE_\gamma} = \frac{1}{4\pi m_\alpha} \cdot \left( \frac{dR_\alpha}{dE_\gamma} \right) \cdot J$$

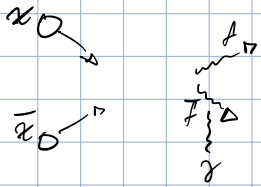
↳ particle physics

J-factor:  $J \equiv \int d\Omega_\psi \int_0^{x_{\text{max}}(\psi)} dx \rho_\alpha[\vec{r}(x)]$

OK, cool, what about annihilation?

$$\psi \cdot \bar{\psi} \longrightarrow f \bar{f} \longrightarrow f \bar{f} + \gamma \gamma \dots$$

( $\psi$  could be its own antiparticle)



if 2 body annihilation:

$$E_i \approx 2m\kappa$$

$$E_f = E_f + E_{\bar{f}} \longrightarrow E_f = E_{\bar{f}} = m\kappa$$

$$|\vec{p}_f| = |\vec{p}_{\bar{f}}|$$

$$E_f = E_{\bar{f}} = m\kappa$$