

BOIZ! let's go!

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Research interests: cosmology stuff, proposing new experiments

11am - 12:20 ERC 545

weekly posts (~50%)

midterm (~20%)

final (~30%)

Foundational topics

Dark Matter Evidence

Speculative Dark Matter Properties

goal: aspects to understand DM research
interdisciplinary
try to avoid theology

What do we know about visible matter?

atoms $\mathcal{O}(\mu\text{m}) \rightarrow$ atom $\mathcal{O}(\text{\AA}) \rightarrow$ protons/neutrons \rightarrow quarks $\mathcal{O}(\text{fm})$
fractional charge

Standard Model of Particle Physics

everything's unstable (will decay into other particles)

"complete" by Higgs Boson

wildly successful

look really good until we look up

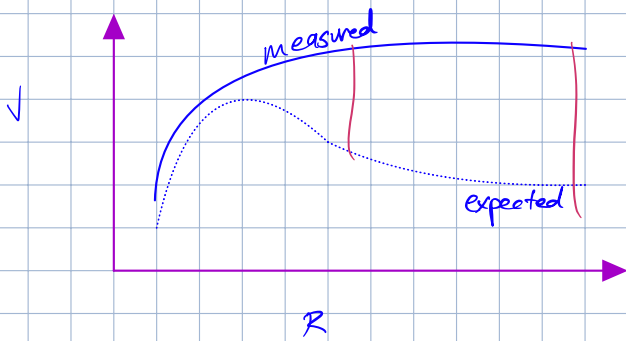
~13.7 billion years since crazy inflation

	13.7 BYA	Today	
Dark Energy	0%	72%	← doesn't dilute
Dark Matter	63%	23%	← like marbles
Atoms	12%	4.6%	
Photons	15%	0%	
Neutrinos	10%	0%	

sensitive to time

What's the evidence

Rotational curves



requires 85% of matter to be "dark"

* should surround
doesn't agree @ large scales

pretty good evidence, not that good

Gravitational lensing

requires ~85% matter to be "dark"

light's affected by gravity

⊗ of curvature to classify mass

Einstein rings

Galaxy Cluster Collisions

~85% total mass passed thru w/o scattering using gravitational lensing

Matter Power Spectrum

how matter clumps

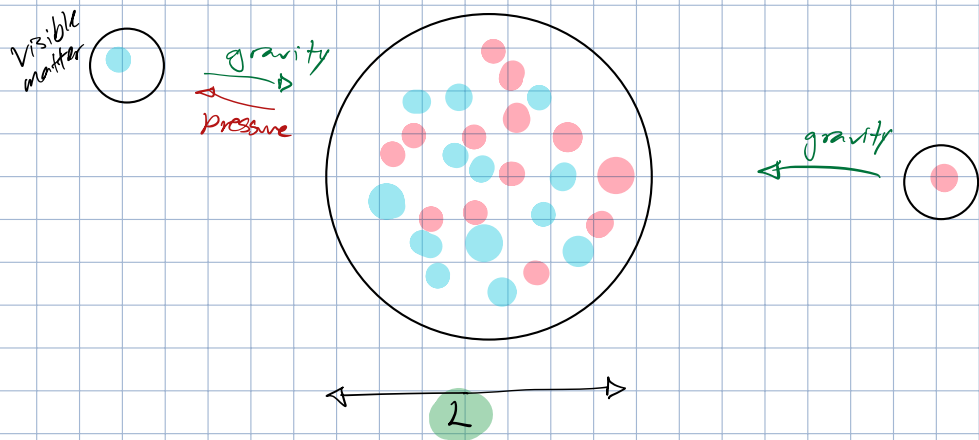
density contrast: $\delta(\vec{x}) = \frac{\rho(\vec{x}) - \bar{\rho}}{\bar{\rho}}$ $\bar{\rho}$ - avg mass density

correlation fn: $\xi(\vec{r}) \equiv \langle \delta(\vec{x}) \cdot \delta(\vec{x} + \vec{r}) \rangle$ fix \vec{r} , then characterize each $\vec{x} + \vec{r}$. we do for different \vec{r} , correlates neighbors

power spectrum: $\xi(\vec{r}) \equiv \int \frac{d^3k}{(2\pi)^3} P(\vec{k}) \cdot e^{i\vec{k} \cdot \vec{r}}$ $k \approx \frac{2\pi}{L}$ **L - length scale**

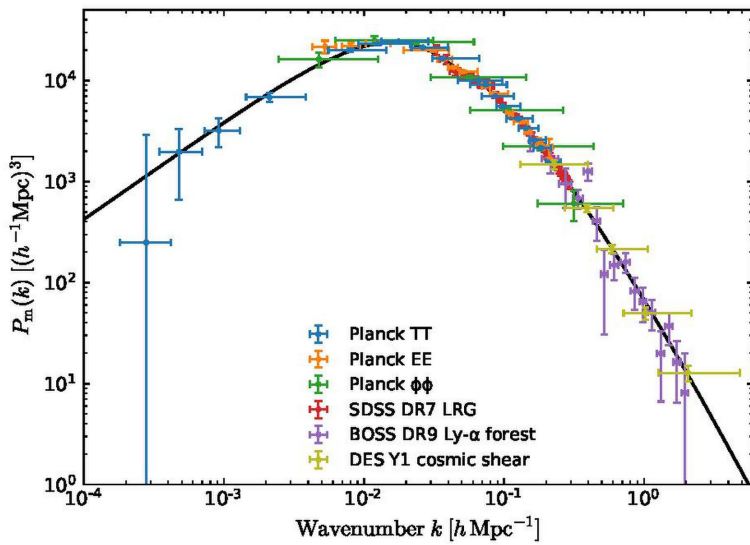
$P(\vec{k})$ is "morally" a 3D Fourier Transform (or waveform decomposition)

overdense region grows by accreting dark matter & visible matter



if $L \gg Mpc$, then we're interested

$P(k)$ is f^2 of L



draw a circle. look @ $S(\vec{x})$ & compare w/neighbors

MOND & $N=DM$ don't agree w/black curve

$$k \equiv \frac{2\pi}{L}$$

CMB

$$\frac{\Delta T}{T} = \sum_{lm} a_{lm} \cdot Y_l^m(\theta, \phi)$$

don't care about density. look @ diff. in temperature

Big Bang Nucleosynthesis

light elements made early on.

can find anti. by measuring today & looking back

missing: how many visible matter needed.

Still! 15% to 85% dark matter

dark matter being made up of visible matter would f*ck things up

Hella evidence @ all scales & time ($z \sim 3400 \rightarrow 0$)

holy grail: extend knowledge to smaller scales

What do we know about DM? (from first principles)

- ① currently based on gravity
- ② extremely stable & long lived
- ③ largely unrelativistic since CMB era
- ④ can't interact much w/ strong or E&M forces
- ⑤ **MIGHT** interact w/ weak or (new) fifth forces

Where to start

SI units

$$[\text{mass}] = \text{kg}$$

$$[\text{temp}] = \text{K}$$

$$[\text{energy}] = \text{J} = \text{kg m}^2 \text{s}^{-2}$$

$$[\text{current}] = \text{A} = \text{C s}^{-1}$$

price to pay: weird constants

$$h = \frac{h}{2\pi} = 1.054 \cdot 10^{-34} \text{ Js}$$

$$c = 2.997 \cdot 10^8 \text{ ms}^{-1}$$

$$R_{\text{atom}} \sim 10^{-10} \text{ cm}$$

$$R_{\text{galaxy}} \sim 10^{23} \text{ cm}$$

much more easy to define

$$c = \hbar = k_B = 1$$

& dimensionless

- ① standard in literature
- ② don't have to keep writing constants
- ③ will learn to love

redefine familiar things

$$c = 2.997 \cdot 10^8 \text{ ms}^{-1} \xrightarrow{\text{seconds}} s = 2.997 \cdot 10^8 \text{ m}$$

$$[\text{length}] = [\text{time}]$$

$$[\hat{x}, \hat{p}] = \hat{x}\hat{p} - \hat{p}\hat{x} = i\hbar$$

$$\text{set } \hbar = 1. \text{ RHS dimensionless} \rightarrow [\text{length}] = [\text{time}] = [\text{momentum}]^{-1}$$

$$E = \frac{mc^2}{\sqrt{1 - \frac{v^2}{c^2}}} \rightarrow \frac{m}{\sqrt{1 - v^2}}$$

$$p = \frac{mv}{\sqrt{1 - \frac{v^2}{c^2}}} \rightarrow \frac{mv}{\sqrt{1 - v^2}}$$

$$[\text{length}] = [\text{time}] = [\text{energy}]^{-1} = [\text{mass}] = [\text{momentum}]^{-1}$$

statistical weights in thermal equilibrium $\exp\left(-\frac{E}{k_B T}\right) \rightarrow \exp\left(-\frac{E}{T}\right)$

$$[\text{length}] = [\text{time}] = [\text{energy}]^{-1} = [\text{mass}] = [\text{momentum}]^{-1} = [\text{temp}]^{-1}$$

electron charge $e = 1.602 \cdot 10^{-19} \text{ C}$

$$e = \sqrt{4\pi\alpha} \quad \alpha = \frac{1}{137}$$

$$c = \frac{1}{\sqrt{\mu_0 \epsilon_0}}$$

$$[\rho] = [\text{volume}]^{-1} = [\text{length}]^{-3}$$

$$\vec{\nabla} \cdot \vec{E} = \rho \rightarrow [\text{E field}] = [\text{length}]^{-2} = [\text{energy}]^{-2}$$

$$\vec{\nabla} \times \vec{B} = \frac{d\vec{E}}{dt} \rightarrow [\text{E field}] = [\text{B field}]$$

$$[\text{E field}] = [\text{B field}] = [\text{energy}]^2$$

standard choice: $[\text{energy}] = \text{GeV} = 10^9 \text{ eV}$

$$[\text{length}] = [\text{time}] = [\text{energy}]^{-1} = [\text{mass}] = [\text{momentum}]^{-1} = [\text{temp}]^{-1} = \text{GeV}$$