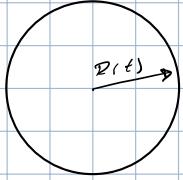


last time



$$\rho(t) = \bar{\rho}(t) (1 + \delta(t))$$

$\delta \ll 1$

$$\begin{aligned} \text{Newton's 2nd Law: } \ddot{R} &= -\frac{G}{R^2} M = -\frac{G}{R^2} \left( \frac{4\pi}{3} R(t)^3 \rho(t) \right) \\ &= -\frac{4\pi}{3} G \cdot R(t) \bar{\rho}(t) (1 + \delta(t)) \end{aligned}$$

$$\text{Mass conservation: } M = \frac{4\pi}{3} R(t)^3 \bar{\rho}(t) (1 + \delta(t)) = \text{const}$$

write  $\frac{\ddot{R}}{R}$  2 different ways  $\rightarrow$  "equation of motion" for density

$$\rightarrow \dot{\delta}(r, t)$$

General EOM for  $\delta(t)$

$$* \ddot{\delta} + 2H\dot{\delta} - \frac{3}{2} \Omega_M H^2 \delta = 0 *$$

<sup>①</sup> Radiation domination ( $T \gg \text{eV}, t \gg 10^6 \text{ yrs}$ )

for simplicity,  $T \gg m_t \approx 175 \text{ GeV}$ ,  $g_*(T) = \text{const} = 106.75$

$$H^2 = \frac{8\pi}{3} G \rho_r = \frac{8\pi}{3} G \left( \frac{T^{2g_*}}{30} T^4 \right) \rightarrow H \propto a^{-2}$$

$$\Omega_M = \frac{\rho_M + \rho_B}{\rho_{tot}} \approx \rho_M$$

$$H = \frac{1}{a} \cdot \frac{da}{dt} = B a^{-2} \rightarrow da/a = B dt \rightarrow \int da/a = B \int dt$$

$$H = \frac{\dot{a}}{a} = \frac{(C_2)t^{-1/2}}{t^{1/2}} = \frac{C_2}{2t}$$

$$\frac{a^2}{2} = B \cdot t \quad a(t) \propto t^{1/2}$$

b/c in RD  $\Omega_M \approx 0$

$$\rightarrow \ddot{\delta} + 2 \left( \frac{1}{2t} \right) \dot{\delta} = 0$$

$$\frac{d^2\delta}{dt^2} = -\frac{1}{t} \cdot \frac{d\delta}{dt} \rightarrow \frac{d^2\delta}{(B/dt)^2} = -\frac{d\delta^2}{t^2}$$

$$\left(\frac{d^2\delta}{dt^2}\right) dt = -\frac{dt^2}{\epsilon}$$

$$d\delta = -\frac{dt}{\epsilon}$$

$$\delta(t) = B + B_2 \log(t)$$

General soln in RD

## ② Matter Domination

$$H^2 = \frac{4\pi}{3} G \rho \underset{\alpha a^{-3}}{\sim}$$

$$H = A \cdot a^{-3/2} = \frac{da}{dt} \frac{1}{a}$$

$$A dt = a^{1/2} da \quad \text{Integrate}$$

$$At = \frac{2}{3} a^{3/2}$$

$$a(t) \propto t^{2/3} \quad \text{in MD}$$

$$H = \frac{\frac{2}{3}t^{-1/3}}{t^{2/3}} = \frac{2}{3t}$$

b/c in MD,  $\Omega_M \approx 1$

$$\ddot{\delta} + 2\left(\frac{2}{3t}\right) \dot{\delta} - \frac{3}{2} \left(\frac{2}{3t}\right)^2 \delta = 0$$

hard. guess polynomial  $\delta(t) = A \cdot t^n \quad \dot{\delta} = A n t^{n-1} \quad \ddot{\delta} = A n(n-1) t^{n-2}$

$$\frac{d^2\delta}{dt^2} + \frac{4}{3t} \frac{d\delta}{dt} - \frac{2}{3t^2} \delta = 0$$

$$A n(n-1) t^{n-2} + \frac{4}{3t} \cdot A n t^{n-1} - \frac{2}{3t^2} \cdot A t^n = 0$$

factor out  $t^{n-2}$

$$A t^{n-2} \left( n(n-1) + \frac{4}{3} n - \frac{2}{3} \right) = 0$$

$$n^2 - n + \frac{4}{3}n - \frac{2}{3} = 0$$

$$n^2 + \frac{1}{3}n - \frac{2}{3} = 0$$

$$n = \frac{-\frac{1}{3} \pm \sqrt{(\frac{1}{3})^2 - 4(1)(-\frac{2}{3})}}{2 \cdot (1)}$$

$$= \frac{-\frac{1}{3} \pm \sqrt{\frac{1}{9} + \frac{8}{3}}}{2} = \frac{-\frac{1}{3} \pm \sqrt{\frac{1}{9} + \frac{24}{9}}}{2}$$

$$= \frac{-\frac{1}{3} \pm \sqrt{\frac{25}{9}}}{2} = \frac{-\frac{1}{3} \pm \frac{5}{3}}{2} = \frac{1}{6}(-1 \pm 5)$$

$$n = \frac{1}{6}(-1 \pm 5)$$

general  $S(t)$  should be linear comb. of both

$$n_1 = \frac{1}{6}(-1+5) = \frac{2}{3} \quad n_2 = \frac{1}{6}(-1-5) = -1$$

$$S(t) = A_1 t^{\frac{2}{3}} + A_2 t^{-1}$$

grow      decay  $\approx 0$

$$\text{in MD, } a(t) \propto t^{\frac{2}{3}}$$

$$\rightarrow S(t) \propto a(t)$$

### ⑤ Dark Energy Domination $t \sim 10 \text{ Gyrs}$

DE doesn't redshift

$$p_1 \propto a^0 \quad \text{NO energy conservation} \rightarrow \text{mass conservation}$$

tho? is local

energy not conserved is global

$$H^2 = \frac{8\pi}{3} G p_1 = \text{const} = \left(\frac{Q^2}{2}\right)^2$$

$$\frac{da}{dt} \frac{1}{a} = Q \rightarrow \frac{da}{at} = Qa \rightarrow a(t) = e^{Qt}$$

$$\frac{\dot{a}}{a} = Q$$

$$\rightarrow \ddot{s} + 2Q\dot{s} = 0 \quad \boxed{\int}$$

$$\dot{s} + 2Qs = 0$$

$$\frac{ds}{dt} = -2Qs \rightarrow s(t) = Ce^{-Qt}$$

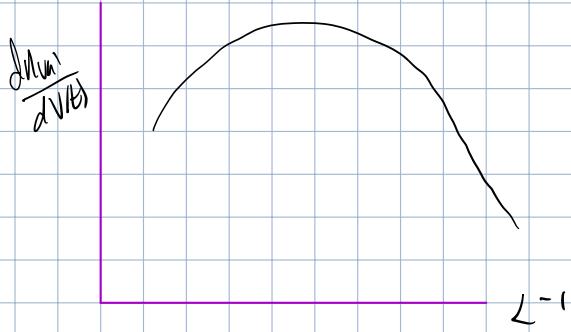
$$\text{perturbations decay} \quad S(t) = Ce^{-Ht}$$

Radiation Domination:  $S(t) = B_2 \ln(t) + B_1$

Matter Domination:  $S(t) = A_1 t^{\frac{2}{3}} \propto a$

Dark Energy Domination:  $S(t) = C e^{-Ht}$

# Matter Power Spectrum



Fourier Transform: generalization of Fourier series

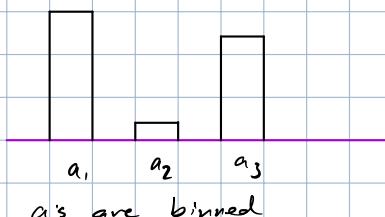
$$\text{Series: } f(x) = \sum a_n \cos\left(\frac{n\pi}{L}x\right) \quad \text{defined in interval } L$$

goal is to determine  $a_n$

$$\text{other form: } f(x) = \sum b_n e^{i \frac{n\pi}{L}x}$$

transform: continuous w/o defined limit

$x$  - pos<sup>n</sup> space  
 $a_n$  - Fourier space



continuous in Fourier transform

2 Point Correlation fn  $C(\vec{R})$

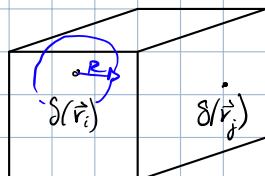
we have  $\delta(\vec{r}, t)$

$$C(\vec{R}) = \left\langle \sum_i \delta(\vec{r}_i) \delta(\vec{r}_i + \vec{R}) \right\rangle$$

$\downarrow$

$\frac{\rho(\vec{r}_i) - \bar{\rho}}{\bar{\rho}}$

angular avg



$$C(\vec{p}) = C(|\vec{R}|)$$